

LSST science reach: two examples Transients and Dark Energy

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May 17, 2017 Sendai Japan

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LSST in one sentence: An optical/near-IR survey of half the sky in ugrizy bands to r~27.5 based on 1000 visits over a 10-year period:

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More information at <u>www.lsst.org</u> and arXiv:0805.2366

A catalog of 17 billion stars and 20 billion galaxies with exquisite photometry, astrometry and image quality!







3.2 Billion Pixel Camera, 10 sq.deg Field





21 science rafts, 189 4K x 4K CCDs



21 "rafts" 9 CCDs per raft

LSST six color system



Data Management must transport, process, archive and serve 15 Tb of raw data / night - 500Pb in 10 years



Algorithm Design





Petascale Computing Design





Data Management Sites and Centers











Main Survey Sky Coverage: number of visits



LSST surveys entire sky south of +5[°] dec with rapid 10 sq.deg exposures

TWO PLANNED SURVEYS:

MAIN SURVEY Deep Wide Survey: 18,000 square degrees to a uniform depth of *u*: 26.1 *g*: 27.4 *r*: 27.5 *i*: 26.8 *z*: 26.1 *y*: 24.9

DEEP DRILLING SURVEY 10% of time: ~20 selected fields. 300 square degrees Continuous 15 sec exposures. 1hour/night

Most of sky covered over 800 times with 30s visits. Alerts on transient objects released worldwide within 60s.

LSST Wide-Fast-Deep survey

A survey of 37 billion objects in space and time

Each sky patch will be visited over 800 times: 30 trillion measurements





- A stream of ~1-10 million time-domain events per night, detected and transmitted to event distribution networks within 60 seconds of observation.
- A catalog of orbits for ~6 million bodies in the Solar System.
- A catalog of ~37 billion objects (20B galaxies, 17B stars), ~7 trillion observations ("sources"), and ~30 trillion measurements ("forced sources"), produced annually, accessible through online databases.
- Deep co-added images.
- Services and computing resources at the Data Access Centers to enable user-specified custom processing and analysis.
- Software and APIs enabling development of analysis codes.

Level

Alert Rate

In ten minutes time the LSST transient pipeline is likely to issue ~80,000 alerts at 5σ .

While most of these will be moving objects, perhaps several thousand will be flaring objects or bursts. Possibly new kinds of objects!

Clearly any followup requires high purity samples. What is needed then is highly trusted event classification. FAST



LSST Services in Support of Level 1



More Information: LSST DPDD; <u>http://ls.st/</u> <u>dpdd</u>

- Forced Photometry ("Precovery") Service
 - Automatically perform and make available forced photometry at the position of all newly detected sources, on imaging acquired over the preceding 30 days
 - On request, for a limited number of positions perform and make available forced photometry on *all* imaging overlapping that position
 - Turnaround on a ~day timescale.
- Alert generation service
 - Package and transmit all alerts to community-supported alert distribution networks
- Limited end-user alert stream subscription and filtering service
 - Individuals will be able to subscribe to receive a small (~20) number of alerts per visit, filtered based on user-specified criteria
 - Limited functionality: no classification, not planning to cross-match to external catalogs, etc.
 - We expect the community will provide <u>sophisticated event brokers</u> with classification engines, cross-match capabilities to other catalogs, etc.

LSST unique contribution: $t < t_0$ information

LSST will have the time history of 37 billion objects in the database. Transient phenomena may exhibit precursor activity:

- 1. Use as Bayesian prior in classification post-alert
- 2. Treat as part of the "light curve." 27th -> 24th mag.
- 3. Catch interesting objects on their way up. Issue pre-alerts.

DARK ENERGY

Measure position and shape of 4 billion galaxies

COSMICTIME

Testing general models of dark energy



Multiple LSST probes of dark energy

- Use the same LSST survey data products
- Analyzed for different signals
- Multiple cross checks
- Combination is far more powerful than root sum of squares

Primary LSST probes	
Weak Lens shear cross correlation tomography Weak Lens magnification tomography	1
2-D Baryon Acoustic Oscillations	√
Supernovae	√
Shear peak statistics	√
Galaxy cluster counts	√
Secondary LSST probes	
Time domain tomography of QSOs and AGNs	\checkmark
Anisotropy of WL+BAO and SN signals	√
New Energy or New Gravity?	

Maximally sensitive to new physics

Integrated Schedule



Science Collaborations



community.Isst.org is a place for the astronomy community to discuss the Large Synoptic Survey Telescope's ongoing development and get help with using LSST's software today. What's here Community members can read, pos... read more

Science-Driven Optimization

of the LSST Observing Strategy

Prepared by the LSST Science Collaborations,

with support from the LSST Project.

Phil Marshall,¹ Scott Anderson,² Timo Anguita,³ Ruth Angus,⁴ Iair Arcavi,⁵ Humna Awan,⁶ Federica B. Bianco,⁷ Rahul Biswas,⁸ Keaton J. Bell,⁹ Eric C. Bellm,¹⁰ David Bennett,¹¹ Niel Brandt, ¹² Chris Britt, ¹³ Derek Buzasi, ¹⁴ Dana I. Casetti-Dinescu, ¹⁵ Laura Chomiuk, ¹⁶ Will Clarkson,¹⁷ Chuck Claver,¹⁸ Andy Connolly,¹⁹ Kem Cook,²⁰ James Davenport,²¹ Victor Debattista,22 Seth Digel,23 Zoheyr Doctor,24 Wen-fai Fong,25 Eric Gawiser,26 Mark Giampapa,27 John E. Gizis, 28 Melissa L. Graham, 29 Carl Grillmair, 30 Zoltan Haiman, 31 Patrick Hartigan, 32 Suzanne Hawley,³³ Željko Ivezić,³⁴ C. Johns-Krull,³⁵ Lynne Jones,³⁶ Shashi Kanbur,³⁷ Vassiliki Kalogera.³⁸ Vinav Kashvap.³⁹ Vishal Kasliwal.⁴⁰ Peter Kurczvnski.⁴¹ Michael C. Liu.⁴² Michelle Lochner,⁴³ Michael B. Lund,⁴⁴ Ashish Mahabal,⁴⁵ Raffaella Margutti,⁴⁶ Tom Matheson,⁴⁷ Peregrine McGehee,48 Søren Meibom,49 Josh Meyers,50 Dave Monet,51 David Nidever,52 Knut Olsen,53 Eric Neilsen,54 Matthew T. Penny,55 Christina Peters,56 Radoslaw Poleski,57 Gordon Richards,³⁸ Stephen Ridgway,⁵⁰ Jeonghee Rho,⁶⁰ Jason Rhodes,⁶¹ David Rubin,⁶² Samuel Schmidt, 63 Ohad Shemmer, 64 Avi Shporer, 65 Colin Slater, 68 Nathan Smith, 67 Marcelles Soares-Santos,68 Keivan Stassun,69 Jay Strader,79 Michael Strauss,71 Rachel Street,72 Christopher Stubbs,⁷³ Paula Szkody,⁷⁴ David Trilling,⁷⁵ Virginia Trimble,⁷⁶ Tony Tyson,⁷⁷ Miguel de Val-Borro,⁷⁸ Stefano Valenti,⁷⁹ Kathy Vivas,⁸⁰ Robert Wagoner,⁸¹ Lucianne Walkowicz,⁸² Beth Willman,⁸³ Peter Yoachim,⁸⁴ Bevin Ashley Zauderer,⁸⁵

Meetings

- working group monthly videocons
- collaboration meetings, hackathons

Community Forums

- public discussions for project & community
- Observing Strategy White Paper
 - science collaborations using OpSim & MAF

Science Roadmaps

Data Challenges

Assessment of Follow-up Capabilities

LSSTC Data Science Fellowship Program - preparing students (hosting opportunities exist)

Science Collaborations

BDESC

Dark Energy Science Collaboration

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Weak Lensing Featured Project: Fast Simulations and Analysis

Submitted by djbard-admin on Fri, 01/30/2015 - 19:48



Caption: Simulated image showing overlapping sources as seen by DES and LSST.

(Author: David Kirkby)

The Weak Lensing Deblending Package provides a framework to produce fast simulations and analysis for weak gravitational lensing.

This software was primarily developed to study the effects of overlapping sources on shear estimation, photometric redshift algorithms, and deblending algorithms. Users can run their own simulations (of LSST and other surveys) or, else download the galaxy catalog and simulation outputs to use with their own code or analyze with the tools provided here.

The code is hosted on github. Please use the github issue tracker to let us know about any issues you

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LSST Project Website project.lsst.org LSST Community Forums community.lsst.org Science Collaborations www.lsstcorporation.org/science-collaborations Observing Strategy White Paper github.com/LSSTScienceCollaborations/ObservingStrategy The 2016 LSST Project and Community Workshop project.lsst.org/meetings/lsst2016 (slides available) Deep Drilling Fields Whitepapers https://project.lsst.org/content/whitepapers32012

Science Drivers to Reference Design (Ivezic et al. 2008; arXiv:0805.2366; Version 4 updated in 2014) LSST Science Requirements (LPM-17) <u>https://docushare.lsstcorp.org/docushare/dsweb/Get/LPM-17</u> LSST Data Products Definitions (LSE-163) <u>https://docushare.lsstcorp.org/docushare/dsweb/Get/LSE-163</u> LSST DM Applications Design (LDM-151) <u>https://docushare.lsstcorp.org/docushare/dsweb/Get/LDM-151</u>

LSST Github Repositories <u>github.com/lsst</u> LSST Software User Guide (The Stack) <u>confluence.lsstcorp.org/display/LSWUG/LSST+Software+User+Guide</u>

LSST Data Science Fellowship Program for Students ciera.northwestern.edu/Education/LSSTC_DSFPOverview.php

Maximizing Science in the Era of LSST: A Community-based Study of Needed US OIR Capabilities <u>noao.edu/meetings/lsst-oir-study</u>

